## Correlated Structure of Nuclear Symmetry Energy from Covariant Nucleon Self-Energy

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## Symmetry Energy in Nuclear Matter

· Equation of state for isospin asymmetric nuclear matter

$$E_b(\rho_b,\delta) = E_0(\rho_b) + E_S(\rho_b)\delta^2 + S_4(\rho_b)\delta^4 + \mathcal{O}(4), \quad L = 3\rho_0 \left. \frac{\partial E_S(\rho_b)}{\partial \rho_b} \right|_{\rho_b = \rho_0},$$

- Important to understand
  - nuclear structure: fission properties, density distribution, collective excitation, ...
  - nuclear astrophysics: NS's radius, crust-core transition density, cooling rate, ...
  - heavy ion reactions: isospin diffusion, DR(n/p), ...



☆ Lattimer:ApJ2013

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*☆ Oertel:RMP2017* 

To improve the nuclear many-body models, information of symmetry energy is essential

New interest in studying ingredients of effective nuclear force: momentum-dependence, exchange term, tensor force, short range correlation, ...

## Symmetry Energy Studied in CDF Theory

Covariant density functional (CDF) theory:

☆ Walecka (1974), Serot (1986), Reihard (1989), Ring (1996),

Bender (2003), Meng (2006), Liang (2015)

- spin-orbit coupling
- pseudo-spin symmetry
- consistent treatment of time-odd fields





## Symmetry Energy Studied in RHF Theory

#### Relativistic Hartree-Fock (RHF) theory:

Bouyssy (1987), Bernardos (1993), Shi (1995),
 Marcos (2004), Long (2006-2010)

- nonlocal Fock terms
- $\pi$ -PV and  $\rho$ -T couplings
- tensor force involved naturally

Improved isospin related structure descriptions # W.H.Long: PLB2006, H.Z.Liang: PRL2008, Q.Zhao: JPG2015



and BYS, PRC 97, 025801 (2018).



→ All mesons account for isospin properties via Fock terms
 → Significant contributions to the symmetry energy from isoscalar meson exchange diagrams

BYS, W.H. Long, J. Meng, and U. Lombardo, PRC 78, 065805 (2008).
 W.H. Long, BYS, K. Hagino, and H. Sagawa, PRC 85, 025806 (2012).
 BYS, Q. Zhao, and W.H. Long, EPJConf 117, 07011 (2016).

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## Nuclear Tensor Interaction: Relativistic Formalism

Relativistic formalism to quantify tensors in Fock diagrams of  $\pi$ -PV,  $\sigma$ -S,  $\omega$ -V,  $\rho$ -T couplings:

L. J. Jiang, S. Yang, BYS, W. H. Long, et al., PRC 91, 034326 (2015).
 Y. Y. Zong and BYS, Chin. Phys. C 42, 024101 (2018).

Second-Order Irreducible Tensor  $S_{12}$  for  $\pi$ -PV:

 $S_{12} = 3 \left( \gamma_0 \boldsymbol{\Sigma}_1 \cdot \boldsymbol{q} \right) \left( \gamma_0 \boldsymbol{\Sigma}_2 \cdot \boldsymbol{q} \right) - \left( \gamma_0 \boldsymbol{\Sigma}_1 \right) \cdot \left( \gamma_0 \boldsymbol{\Sigma}_2 \right) \boldsymbol{q}^2$ 

 $\rightarrow$  The tensors are involved naturally by the Fock diagrams and quantified by the relativistic formalism without introducing any additional free parameters.

Two evidence: Spin dependence, Tensor Sum Rule





0.6

# Q. Zhao, BYS, W. H. Long, J. Phys. G 42, 095101 (2015).

		TW99	PKDD	PKO1	PKA1	BHF
J	T = 0	51.0	50.8	38.8	42.4	44.2
	T = 1	-26.2	-22.1	-8.1	-5.7	-9.0
	kin	8.0	8.1	3.7	0.5	-1.0

BHF: & I. Vidaña+:PRC2011 CFG: & Or Hen+:PRC2015

0.4 PKA1 0.2 σ-S 0.0 **a**-- 1n -0.2 -- 0-- 1d 0.6 Tensor AE<sub>so</sub> (MeV) 0.4 02 0.0 -0.2 0.6 H<sup>T,</sup> & H Centra 0.4 0.2 0.0 -0.2 -0.4 1p3/2 1p<sub>1/2</sub> 1p.,, 1d,... 1d\_... 1d,,,, 1p.,, 1d... Neutron (v) orbit j'

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DDRHF

## Impact of Tensor Force on Nuclear Symmetry Energy

Tensor force effects on kinetic symmetry energy:									
	TW99 PKDD PKO1 PKA1								
	$E_S$	kin	8.0	8.1	3.7	0.5			
		kin-T <sub>1</sub>			-7.3	-9.7			
		kin- $T_2$			-0.4	-0.6			

☆ Or Hen et al., PRC 91, 025803 (2015). -10±7.5 MeV
 ☆ B. J. Cai, B. A. Li, PRC 92, 011601 (2015). -14.28±11.60 MeV
 ☆ B. J. Cai, B. A. Li, PRC 93, 014619 (2016). -16.94±13.66 MeV

Effects of Fock terms to  $E_S$ : Soften due to tensor part



*☆ L. J. Jiang et al., PRC* **91**, 025802 (2015).



Nuclear fourth-order symmetry energy:  $S_4$  suppressed in RHF, but  $S_{4,kin}$  enhanced at  $\rho_0$  $\Rightarrow Z.W. Liu et al., PRC$ **97**, 025801 (2018).

		TW99	PKDD	PKO1	PKA1
$S_4$	kin	0.5	0.5	1.2	2.1
	kin-T			0.2	0.3

0.0 0.2 0.4 0.6 0.8 0.2 0.4 0.6 0.8 0.2 0.4 0.6 0.8  $\Leftrightarrow B. J. Cai, B. A. Li, PRC$  92, 011601 (2015). 7.18 ± 2.52 MeV  $\rho$  (fm<sup>3</sup>)

## Properties of $E_S$ at Saturation Density: J and L

Methods to analyse the structure of symmetry energy:

- In terms of energy density functional: kinetic and potential, spin-isospin
- In terms of components of nuclear force: central and tensor part
- In terms of single-particle energy: nucleon self-energy

## Properties of $E_S$ at Saturation Density: J and L

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Properties of Symmetry Energy at Saturation Density in CDF theory:

From a covariant nucleon self-energy, try to understand:

- Origin of model dependence of L
- Correlations between L and J

☆ Z. W. Liu, Q. Zhao and BYS, arXiv:1809.03837.

## Hugenholtz-Van Hove (HVH) Theorem

- Symmetry Energy from HVH theorem

☆ B. J. Cai, L. W. Chen, Phys. Lett. B 711 (2012) 104.

$$E_{b} + \rho_{b} \frac{\partial E_{b}}{\partial \rho_{b}} = \varepsilon_{F} \xrightarrow{expansion}_{io asymmetry} E_{S}(\rho_{b}) = \frac{1}{4} \frac{d}{d\delta} \Big[ \sum_{\tau} \tau \varepsilon_{F}^{\tau}(\rho_{b}, \delta, k_{F}^{\tau}) \Big] \Big|_{\delta=0}$$
$$\frac{d\varepsilon_{F}^{\tau}}{d\delta} |_{\delta=0} = \frac{\tau k_{F}}{3} \Big( \frac{\partial \varepsilon}{\partial k} \Big) \Big|_{k=k_{F}} + \frac{\partial \varepsilon_{F}^{\tau}}{\partial \delta} \Big|_{\delta=0} \implies E_{S}(\rho_{b}) = \boxed{E_{S}^{kin}(\rho_{b}) + E_{S}^{mon}(\rho_{b})} + \boxed{E_{S}^{1st}(\rho_{b})}$$

• Kinetic part, k-dependence of self-energy,  $\delta$ -dependence of self-energy

$$\begin{split} E_{S}^{\mathrm{kin}}(\rho_{b}) &= \frac{k_{F}k_{F}^{*}}{6E_{F}^{*}}, \qquad E_{S}^{\mathrm{1st}}(\rho_{b}) = \frac{1}{2}\frac{g_{\rho}^{2}}{m_{\rho}^{2}}\rho_{b} + E_{S}^{\mathrm{1st},E} \\ E_{S}^{\mathrm{mon}}(\rho_{b}) &= \frac{k_{F}k_{F}^{*}}{6E_{F}^{*}}\frac{\partial\Sigma_{V}}{\partial k}\big|_{k=k_{F}} + \frac{k_{F}M_{F}^{*}}{6E_{F}^{*}}\frac{\partial\Sigma_{S}}{\partial k}\big|_{k=k_{F}} + \frac{k_{F}}{6}\frac{\partial\Sigma_{0}}{\partial k}\big|_{k=k_{F}}, \end{split}$$

• Symmetry energy at saturation density

$$J = J^{kin} + J^{mom} + J^{1st}$$
$$L = L^{kin} + L^{mom} + L^{1st} + L^{cross} + L^{2nd}$$

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$$J = J^{\text{kin}} + J^{\text{mom}} + J^{1\text{st}}, \qquad L = L^{\text{kin}} + L^{\text{mom}} + L^{1\text{st}} + L^{\text{cross}} + L^{2\text{nd}}$$

$$\int_{100}^{100} \int_{100}^{100} \int_{100}^{10$$

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 $\varepsilon^{*,\tau} = \varepsilon^{\tau} - \Sigma_0^{\tau}(\rho, \delta, |\mathbf{k}|),$ 

$$J = J^{\text{kin}} + J^{\text{mom}} + J^{1\text{st}}, \qquad L = L^{\text{kin}} + L^{\text{mom}} + L^{1\text{st}} + L^{\text{cross}} + L^{2\text{nd}}$$



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$$\int_{0}^{78} \int_{0}^{66} \int_{0$$

☆ Z. W. Liu, Q. Zhao and BYS, arXiv:1809.03837.

Two underlying linear correlations construct the fundamental correlation between *L* and *J* in CDF framework.

•  $\Sigma_{\mathcal{O}}^{\text{sym},1}$ : isovector scalar coupling

•  $\Sigma_{\mathcal{O}}^{\text{sym},2}$ : strong model dependence

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## Landau Mass in CDF Theory



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## Neutrino Emissivity in DUrca Process

• Non-relativistic expression:

*☆ J.M. Lattimer et al., PRL* **66** (2017) 2701.

$$Q^{(D)} = \frac{457\pi}{10080} G_F^2 \cos^2 \theta_c (F_V^2 + 3G_A^2) \frac{m_n^* m_p^* m_l^*}{\hbar^{10} c^3} (k_B T)^6 \Theta_{np}$$

Adopting a non-relativistic approximation of the matrix element.

$$\begin{aligned} \mathcal{Q}^{(D)} &= \frac{457\pi}{10080} \frac{G_F^2 \cos^2 \theta_c}{\hbar^{10} c^3} (k_B T)^6 [F_V G_A((\varepsilon_{F_n} + \varepsilon_{F_p}) p_{Fl}^2 - (\varepsilon_{F_n} - \varepsilon_{F_p}) (p_{F_n}^2 - p_{F_p}^2)) + 2G_A^2 \mu_l M_n^* M_p^* \\ &+ (F_V^2 + G_A^2) (\mu_l (2\varepsilon_{F_n} \varepsilon_{F_p} - M_n^* M_p^*) + \varepsilon_{F_n} p_{Fl}^2) - \frac{1}{2} (\varepsilon_{F_n} + \varepsilon_{F_p}) (p_{F_n}^2 - p_{F_p}^2 + p_{Fl}^2)] \Theta_{npe} \end{aligned}$$

Relativistic matrix element, but approximation in phase space integration

• New expression for neutrino emissivity in DUrca processes:

$$\begin{split} \mathcal{Q}^{(D)} &= \mathcal{Q}_{1}^{(D)} + \mathcal{Q}_{2}^{(D)} \\ \mathcal{Q}_{1}^{(D)} &= \frac{457\pi}{10080} \frac{G_{F}^{2} \cos^{2} \theta_{c}}{\hbar^{10} c^{3}} (k_{B}T)^{6} [2(F_{V}^{2} + G_{A}^{2}) + \hat{M}_{n} \hat{M}_{p} (G_{A}^{2} - F_{V}^{2})] m_{n}^{*} m_{p}^{*} m_{l}^{*} \Theta_{npe} \\ \mathcal{Q}_{2}^{(D)} &= -\frac{45[2\pi^{2} \zeta(5) + 63\zeta(7)]}{16\pi^{5}} G_{F}^{2} \cos^{2} \theta_{c} \frac{m_{n}^{*} m_{p}^{*}}{\hbar^{10} c^{5}} \hat{P}_{n} \hat{P}_{p} \left[ \frac{|\mathbf{p}_{nF}|}{|\mathbf{p}_{pF}|} (F_{V} + G_{A})^{2} + \frac{|\mathbf{p}_{pF}|}{|\mathbf{p}_{nF}|} (F_{V} - G_{A})^{2} \right] (k_{B}T)^{7} \Theta_{npe} \end{split}$$

Relativistic matrix element, more attention to angular integration in phase space

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## Neutrino Emissivity in DUrca Process



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Relativistic matrix element, more attention to angular integration in phase space

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## Summary and Outlook

- The symmetry energy in CDF theory, illustrating the effects of exchange terms, also the tensor force effects in kinetic part, impact on Landau mass and neutrino emissivity
- The correlated structure between *L* and *J* is revealed in terms of the nucleon self-energy
- Possible way to improve the CDFs in the market, from a viewpoint of nucleon self-energy, constrained by either ab initios or experiments



Still on the way to establish an unified CDF both for finite nuclei and for astrophysics

## Summary and Outlook

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# Thank you for your attention!

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## Selected CDF Effective Lagrangians

Table: Bulk properties of symmetric nuclear matter at saturation point

	Fock	$\sigma$ -NL	$\omega\text{-NL}$	DD	$\pi$ -PV	$\rho$ -T	$\rho_0$	$E_B/A$	Κ	J	L	Reference
							$(fm^{-3})$	(MeV)	(MeV)	(MeV)	(MeV)	
PKA1	<b>√</b>	×	×	$\checkmark$	$\checkmark$	$\checkmark$	0.160	-15.83	230.0	36.0	104	Long:2007
PKO1	$\checkmark$	×	×	$\checkmark$	$\checkmark$	×	0.152	-16.00	250.2	34.4	98	Long:2006
PKO2	$\checkmark$	×	×	$\checkmark$	×	×	0.151	-16.03	249.6	32.5	76	Long:2008
PKO3	$\checkmark$	×	×	$\checkmark$	$\checkmark$	×	0.153	-16.04	262.5	33.0	83	Long:2008
NL1	×	$\checkmark$	×	×	×	×	0.152	-16.43	211.2	43.5	140	Reinhard:1986
NL3	×	$\checkmark$	×	×	×	×	0.148	-16.25	271.7	37.4	118	Lalazissis:1997
NL-SH	×	$\checkmark$	×	×	×	×	0.146	-16.33	354.9	36.1	114	Sharma:1993
TM1	×	$\checkmark$	$\checkmark$	×	×	×	0.145	-16.26	281.2	36.9	111	Sugahara:1994
PK1	×	$\checkmark$	$\checkmark$	×	×	×	0.148	-16.27	282.7	37.6	116	Long:2004
TW99	×	×	×	$\checkmark$	×	×	0.153	-16.25	240.3	32.8	55	Typel:1999
DD-ME1	×	×	×	$\checkmark$	×	×	0.152	-16.20	244.7	33.1	56	Nikšić:2002
DD-ME2	×	×	×	$\checkmark$	×	×	0.152	-16.11	250.3	32.3	51	Lalazissis:2005
PKDD	×	×	×	$\checkmark$	×	×	0.150	-16.27	262.2	36.8	90	Long:2004

Relatively large values of K and J systematically in RMF with nonlinear self-coupling of mesons (NLRMF)

☆ B. Y. Sun et al., PRC 78(2008)065805; W. H. Long et al., PRC 85(2012)025806; L. J. Jiang et al., PRC 91(2015)025802.

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#### **Relativistic Formalism of Tensors**

**Relativistic formalism** to quantify tensors in Fock diagrams of  $\pi$ -PV,  $\sigma$ -S,  $\omega$ -V,  $\rho$ -T couplings:  $\Leftrightarrow$  L. J. Jiang, S. Yang, B. Y. Sun, W. H. Long, and H. Q. Gu, Phys. Rev. C **91**, 034326 (2015).

$$\mathcal{H}_{\pi-\mathrm{PV}}^{T} = -\frac{1}{2} \left[ \frac{f_{\pi}}{m_{\pi}} \bar{\psi} \gamma_0 \Sigma_{\mu} \vec{\tau} \psi \right]_1 \cdot \left[ \frac{f_{\pi}}{m_{\pi}} \bar{\psi} \gamma_0 \Sigma_{\nu} \vec{\tau} \psi \right]_2 D_{\pi-\mathrm{PV}}^{T, \ \mu\nu}(1, 2), \tag{2}$$

$$\mathcal{H}_{\sigma-\mathrm{S}}^{T} = -\frac{1}{4} \Big[ \frac{g_{\sigma}}{m_{\sigma}} \bar{\psi} \gamma_{0} \Sigma_{\mu} \psi \Big]_{1} \Big[ \frac{g_{\sigma}}{m_{\sigma}} \bar{\psi} \gamma_{0} \Sigma_{\nu} \psi \Big]_{2} D_{\sigma-\mathrm{S}}^{T, \ \mu\nu}(1, 2), \tag{3}$$

$$\mathcal{H}_{\omega-\mathrm{V}}^{T} = +\frac{1}{4} \Big[ \frac{g_{\omega}}{m_{\omega}} \bar{\psi} \gamma_{\lambda} \gamma_{0} \Sigma_{\mu} \psi \Big]_{1} \Big[ \frac{g_{\omega}}{m_{\omega}} \bar{\psi} \gamma_{\delta} \gamma_{0} \Sigma_{\nu} \psi \Big]_{2} D_{\omega-\mathrm{V}}^{T, \ \mu\nu\lambda\delta}(1,2), \tag{4}$$

$$\mathcal{H}_{\rho,\mathrm{T}}^{T} = +\frac{1}{2} \left[ \frac{f_{\rho}}{2M} \bar{\psi} \sigma_{\lambda\mu} \vec{\tau} \psi \right]_{\mathrm{I}} \cdot \left[ \frac{f_{\rho}}{2M} \bar{\psi} \sigma_{\delta\nu} \vec{\tau} \psi \right]_{2} D_{\rho,\mathrm{T}}^{T, \ \mu\nu\lambda\delta}(1,2), \tag{5}$$

where  $\Sigma^{\mu} = (\gamma^5, \Sigma)$ , and  $D^T$  ( $\phi$  for  $\sigma$  and  $\pi$ ,  $\phi'$  for  $\omega$  and  $\rho$ ) read as,

$$D_{\phi}^{T, \ \mu\nu}(1,2) = \left[\partial^{\mu}(1)\partial^{\nu}(2) - \frac{1}{3}g^{\mu\nu}m_{\phi}^{2}\right]D_{\phi}(1,2) + \frac{1}{3}g^{\mu\nu}\delta(x_{1} - x_{2}),$$
  
$$D_{\phi'}^{T, \ \mu\nu\lambda\delta}(1,2) = \left[\partial^{\mu}(1)\partial^{\nu}(2)g^{\lambda\delta} - \frac{1}{3}G^{\mu\nu\lambda\delta}m_{\phi'}^{2}\right]D_{\phi'}(1,2) + \frac{1}{3}G^{\mu\nu\lambda\delta}\delta(x_{1} - x_{2}).$$

$$G^{\mu\nu\lambda\delta} \equiv \left(g^{\mu\nu}g^{\lambda\delta} - \frac{1}{3}g^{\mu\lambda}g^{\nu\delta}\right)$$

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#### Isospin and Tensor Effects on Symmetry Energy

$$E_{S}(\rho_{b}) = E_{S,k} + E_{S,T=0}^{D} + E_{S,T=0}^{E} + E_{S,T=1}^{D} + E_{S,T=1}^{E}$$



		TW99	PKDD	PKO1	PKA1	BHF
	kin	5.9	5.0	-34.5	-69.6	14.9
L	T = 0	62.2	78.2	67.5	71.3	69.1
	T = 1	-12.8	7.0	64.8	103.2	-17.5

Large model dependence in kinetic and T=1 potential parts: Significant  $E_S^{E,\sigma+\omega}$ 

Stiff  $E_s$  in RHF: Too strong T=1 $\bigcirc$ 

*☆ I. Vidaña et al., PRC* 84, 062801(*R*) (2011).

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## Effective Mass in RHF Theory

• Definition of the Effective Mass:

$$\frac{M_{\tau}^{*}}{M} = 1 - \frac{dU_{\tau}(k,\epsilon(k))}{d\epsilon_{\tau}} = k \frac{dk}{d\epsilon_{\tau}} = \left[1 + \frac{M}{k} \frac{dU_{\tau}}{dk}\right]^{-1}$$

• Non-relativistic Mass: Shrödinger equivalent potential  $U_e^{\tau}$ 

$$E - mass: \frac{\bar{M}}{M} = \left[1 - \frac{\partial U^{\tau}}{\partial \epsilon^{\tau}}\right], \quad K - mass: \frac{\tilde{M}}{M} = \left[1 + \frac{M}{k} \frac{\partial U^{\tau}}{\partial k}\right]^{-1}$$

$$\frac{M_{NR}^*}{M} = \frac{\bar{M}}{M} \cdot \frac{\tilde{M}}{M} = \left[1 - \frac{\Sigma_0}{M}\right] \left[1 + \left(\frac{M^*}{k}\frac{\partial\Sigma_S}{\partial k} + \frac{E^*}{k}\frac{\partial\Sigma_0}{\partial k} + \frac{k^*}{k}\frac{\partial\Sigma_V}{\partial k} + \frac{\Sigma_V}{k}\right)\right]^{-1}$$

• Landau Mass:  $\epsilon + M = E^* + \Sigma_0$ 

$$M_L^* = k \frac{dk}{d\epsilon^{\tau}} = k \Big[ \frac{k^*}{E^*} + \frac{M^*}{E^*} \frac{\partial \Sigma_{S,E}}{\partial k} + \frac{k^*}{E^*} \frac{\partial \Sigma_{V,E}}{\partial k} + \frac{\partial \Sigma_{0,E}}{\partial k} \Big]^{-1}$$

• Relativistic Mass:

☆ W. H. Long et al., Phys. Lett. B 640, 150 (2006).

$$\frac{M_R^*}{M} = 1 - \frac{d}{d\epsilon} \Big[ U_e^{\tau} - \frac{\epsilon^2}{2M} \Big], \quad M_R^* = M_{NR}^* + \epsilon = M_g^* \rightarrow \text{group mass}$$

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